

The Galactic Scale Effects of Active Galactic Nuclei

Fabry-Pérot observations of NGC 1068

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Obscured AGN: Insight into AGN Evolution







Goulding et al. 2014

Mitra et al. 2018

RSS: An Effective Tool to Study Type II AGN

- Study emission line ratios of large samples of AGN. (e.g. Hainline et al. 2014; Hviding et al. 2018)
- Used to characterize a bright subset of the extremely obscured quasar population.



SALT Queue based observing ideal for this study as we primarily wanted to find redshift and strong AGN lines, which can be done in a variety of conditions

RSS Fabry-Pérot on SALT

Advantages:

Large FOV (4' Radius)

Calibrated for use in:

- 480 515 nm,
- 530 600 nm,
- 650 705 nm,
- 730 770 nm

LR mode can resolve nearby lines, e.g. [NII] from Hα Challenges & Opportunities:

The wavelength solution is difficult to constrain without strong sky features

The available pipeline is incomplete and no longer maintained



NGC 1068 Fabry-Pérot Results

Velocity Field



NGC 1068 Fabry-Pérot Results

Emission Line Intensity Maps



Largest FOV resolved H α and [NII] maps of NGC 1068

Line Ratios as an AGN diagnostic

Traditionally, AGN are classified using the Baldwin, Phillips & Terlevich (1981) Diagram



NGC 1068 Fabry-Pérot Results



Hviding et al. in prep.

Longslit Confirmation



Longslit observations to constrain BPT y-axis ($log_{10}([OIII]/H\beta)$)

TYPHOON PrISM Survey Results

- Progressive Integral
 Step Method (PrISM)
 Observations of the 100
 closest and largest
 galaxies in the Southern
 Hemisphere
- 2.5m duPont telescope at Las Campanas Observatory
- These are first results, further details are forthcoming.



TYPHOON PrISM Survey Results

The same ionization features are observed



D'Agostino et al. 2018

Ionization Features as Outflows

Evidence:

- High ionization parameters in both features, consistent with an AGN from the BPT
- D'Agostino et al. 2018 argue that the features are consistent with radio jets/lobes
- Not to scale, about 30X smaller than shown



Ionization Features as Outflows

Problems



An Alternate Interpretation: Light Echoes

Light echoes have been observed in AGN to characterize fading AGN candidates



Keel et al. 2016

AGN light echoes may shed light into the characteristic timescales of mass accretion and AGN evolution



Yang et al. 2018

An Alternate Interpretation: Light Echoes

De-projecting and de-lagging NGC 1068



An Alternate Interpretation: Light Echoes



Hviding et al. in prep.

Conclusions & Questions

Conclusions:

RSS Fabry-Pérot successful for studying large nearby sources

Ionization features in NGC 1068 may be evidence for light echoes rather than outflows.

The ionization features will shed light on AGN evolution and the role of Type II sources

Opportunities for Future Work

Completing and maintaining the existing FP pipeline with SALT

Distinguishing between shock ionization and photoionization

Quantifying the inclination of NGC 1068 to constrain de-lagged physical distances to nucleus